

IceDUNE: oscillations and friends!

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Imagine what can happen if they partner up...





Some thoughts on what else could be done with it



Redundancy = robustness

How to be redundant in the the neutrino sector?







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Leading measurements of CP violation will be done with accelerator neutrinos





Sub-GeV atmospheric neutrinos are one of the richest neutrino samples we have access to.







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But sub-GeV atmospherics are very difficult...



Needs to know neutrino direction

Low E protons are invisible

② Cherenkov detectors

Liquid Argon TPCs can do it!















ArgoNeuT demonstrated the LAr capability to detect 21 MeV recoil protons. AraoNeuT 1810.06502







LArIAT 1911.10379

Muons:

$$\mu^+ \rightarrow e^+ \overline{v}_{\mu} v_e$$

 $\mu^- p^+ \rightarrow v_{\mu} + n$

Pions:

 $\pi^+ n \rightarrow \pi^0 p^+$ $\pi^- p^+ \rightarrow \pi^0 n$

Topology depends on particle and its charge





Simulate neutrino-argon interactions with event generators

Use realistic atmospheric fluxes (Honda et al 1502.03916)

Account for uncertainties of atmospheric neutrino fluxes 40% normalization, 5% e/ μ ratio, 2% nu/nubar ratio, ± 0.2 spectral distortion coefficient

Realistic LArTPC capabilities $\Delta p = 5\%$, 5%, 10%, $\Delta \theta = 5^{\circ}$, 5°, 10°, for e, µ, p, K_p = 30 MeV

Classify events by final state topology (number of protons)

 $\Phi_{\alpha}(E) = \Phi_{\alpha,0} f_{\alpha}(E) \left(\frac{E}{E_0}\right)'$







1.									1	1
	-10.63	10.46	10.41	10.53	11.91	11.75	10.20	9.30	8.66	5.94 -
).8	- -23.04 	22.69	22.31	22.44	25.16	24.81	21.63	19.84	18.62	- 12.52-
	-31.69	31.59	30.45	30.33	33.60	33.07	29.00	26.82	25.54	16.46-
n e	-41.31	41.84	39.49	39.00	42.43	41.77	36.90	34.34	33.23	20.66-
5.0	-52.43	54.38	50.42	49.60	52.97	52.07	46.32	43.20	42.54	26.00-
ר ר	-64.67	68.96	63.06	61.65	64.31	63.25	56.80	53.05	53.14	31.83-
J.4	-73.42	80.14	72.59	70.60	71.94	70.66	64.06	60.11	60.78	35.21-
חח	-60.29	72.12	65.94	63.72	63.50	62.15	56.65	53.34	53.75	30.03-
J.Z	-26.49	33.30	30.80	30.21	29.71	28.85	26.09	24.24	23.90	13.09-
_	1().8-().6-().4-0	0.2 0). 0.	2 0.	4 0.	6 0.	8 1
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 N_e – CC–1p0 $\pi,\,\delta_{\rm CP}$ = $3\pi/2$

 $\Delta N_e - CC - 1p0\pi$, $\delta_{CP} = 3\pi/4$

										-
	1.	0.65	-0.48	-0.21	-0.90	-0.58	-0.08	0.00	0.00	
Ear [GeV]	0.8	1.45	-1.07	-0.55	-1.89	-1.26	-0.22	-0.02	0.00	
		2.31	-1.54	-0.90	-2.67	-1.75	-0.40	-0.05	-0.01	
	0.6		-2.26	-1.46	-3.58	-2.55	-0.72	-0.13	-0.03	
	0.0	4.74	-3.45	-2.31	-4.66	-3.39	-1.16	-0.28	-0.08	
	0.4	-6.28	-4.89	-3.39	-6.01	-4.86	-2.10	-0.67	-0.21	
	0.4		-7.19	-4.69	-6.84	-6.09	-3.25	-1.39	-0.54	
	0.2	-6.14	-8.55	-4.71	-5.54	-5.33	-3.44	-1.88	-0.92	
	0.2	1.08	-2.93	-2.46	-2.54	-2.53	-1.88	-1.16	-0.66	
	_	1().8-0).6-().4-(0.2 0). 0.	2 0.	.4 0.	•
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 $\cos \theta_z^{\text{dep}}$











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DUNE has a unique opportunity to study sub-GeV neutrinos

This opens up the possibility of measuring CP violation independently of the beam







Some thoughts on what else could be done with it





DUNE: MSW resonance at the solar splitting

What do we learn when we combine both?

IceCube: MSW resonance at the atmospheric splitting



NSIs are relevant for high energy, but also low energy due to solar Δm^2 driven oscillations





Sterile neutrinos

Idea: Nunokawa, Peres, Zukanovich-Funchal hep-ph/0302039





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Sterile neutrinos









Sterile Oscillations at High and Low Energies







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LV	'nT	2021







Conclusions

- Atmospheric neutrinos still have lots of potential
 - Low energy, high energy
 - CP violation, steriles, NSIs, novel particles, ...
 - We need to get to work.



Backup







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